Simulating all-scale global weather with the Finite-Volume Module of IFS

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$$\begin{split} &\frac{\partial\mathcal{G}\rho}{\partial t}+\nabla\cdot(\mathbf{v}\mathcal{G}\rho)=0\\ &\frac{\partial\mathcal{G}\rho\mathbf{u}}{\partial t}+\nabla\cdot(\mathbf{v}\mathcal{G}\rho\mathbf{u})=\mathcal{G}\rho\left(-\Theta_d\bar{\mathbf{G}}\nabla\varphi'-\frac{\mathbf{g}}{\theta_a}\left(\theta'+\theta_a(\mathbf{e}_d'-\mathbf{v}_a,\psi)+\mathbf{f}'\cdot(\mathbf{u}'-\frac{\mathbf{g}}{\theta_a}\mathbf{u})+\mathbf{M}(\mathbf{u})+\mathbf{D}\right)\\ &\frac{\partial\mathcal{G}\rho\mathbf{u}}{\partial t}+\nabla\cdot(\mathbf{v}\mathcal{G}\rho\theta')=\mathcal{G}\rho\left(-\bar{\mathbf{G}}^T\mathbf{u}\cdot\nabla\theta_a-\frac{L}{c_p\pi}\left(\frac{\Delta q_{ex}}{\Delta t}+\mathbf{b}\right)+\mathbf{h}\right)\\ &\frac{\partial\mathcal{G}\rho\mathbf{g}}{\partial t}+\nabla\cdot(\mathbf{v}\mathcal{G}\rho\mathbf{g})=\mathcal{G}\rho\mathcal{R}^{h_b}\\ &\frac{\partial\mathcal{G}\rho\varphi'}{\partial t}+\nabla\cdot(\mathbf{v}\mathcal{G}\rho\varphi')=\mathcal{G}\rho\sum_{\ell=1}^3\left(\frac{a_\ell}{\zeta_\ell}\nabla\cdot\zeta_\ell(\bar{\mathbf{v}}-\bar{\mathbf{G}}^T\mathbf{C}\nabla\varphi')\right)+b_\ell, \end{split}$$



Operational configuration of the Integrated Forecasting System at ECMWF



Current operational configuration of the Integrated Forecasting System (IFS) at the European Centre for Medium-Range Weather Forecasting:

- hydrostatic primitive equations (nonhydrostatic option available; see Benard et al. 2014)
- hybrid η p vertical coordinate (Simmons and Burridge, 1982)
- spherical harmonics representation in horizontal (Wedi et al., 2013)
- finite-element discretisation in vertical (Untch and Hortal, 2004)
- semi-implicit semi-Lagrangian (SISL) integration scheme (Temperton et al. 2001, Diamantakis 2014)
- cubic-octahedral ("TCo") grid (Wedi, 2014, Malardel et al. 2016)
- HRES: TCo1279 (O1280) with $\Delta_h \approx$ 9 km and 137 vertical levels
- ENS (1+50 perturbed members): TCo639 (O640) with $\Delta_h \approx 16$ km and 91 vertical levels
- \Rightarrow ECMWF strategy for the year 2025 targets to run ENS with TCo1999 with $\Delta_h \approx 5$ km

Greyzone evaluations



Total water + ice content

tqc 2016081100 +13h





Quasi-hydrostatic versus nonhydrostatic dynamics

Idealized convective storm (Klemp et al. 2015) on a small planet (1/25 reduced) with H and NH formulation of IFS: From what horizontal grid spacing Δ_h appear significant differences?



 \rightarrow H-IFS and NH-IFS use Forbes et al. 2011 microphysics and similar numerical configurations, in particlar TCo grid, FD in vertical, ICI, no explicit diffusion, no convection scheme)

Quasi-hydrostatic versus nonhydrostatic dynamics

Idealized convective storm (Klemp et al. 2015) on a small planet (1/25 reduced) with H and NH formulation of IFS and NH-FVM:



 \rightarrow NH-FVM uses smaller time steps and different microphysics parametrisation!



Finite-Volume Module (FVM) of IFS-key formulation features

- deep-atmosphere nonhydrostatic Euler equations in geospherical framework (Szmelter and Smolarkiewicz 2010; Smolarkiewicz et al. 2016; Smolarkiewicz, Kühnlein, Grabowski 2017; Kühnlein, Malardel, Smolarkiewicz in prep.)
- · flexible height-based terrain-following vertical coordinate
- hybrid of horizontally-unstructured median-dual finite-volume with vertically-structured finite-difference/finite-volume discretisation (Szmelter and Smolarkiewicz 2010; Smolarkiewicz et al. 2016)
- · all prognostic variables are co-located
- two-time-level semi-implicit integration scheme with 3D implicit acoustic, buoyant and rotational modes (Smolarkiewicz, Kühnlein, Wedi JCP 2014)
- preconditioned generalised conjugate residual iterative solver for 3D elliptic problems arising in the semi-implicit integration schemes (Smolarkiewicz and Szmelter 2011 for a more recent review)
- non-oscillatory finite-volume MPDATA scheme (Smolarkiewicz and Szmelter 2005; Kühnlein and Smolarkiewicz 2017)
- octahedral reduced Gaussian grid, but the FVM formulation not restricted to this (Szmelter and Smolarkiewicz 2016)
- optional moving mesh capability (as in Kühnlein, Smolarkiewicz, Dörnbrack 2012)





dual volume: V_i , face area: S_i



median-dual finite-volume approach

Octahedral reduced Gaussian grid



Nodes of octahedral grid 'O24'



Primary mesh about nodes of octahedral grid in FVM



- octahedral reduced Gaussian grid (octahedral grid of size OX)
- suitable for spherical harmonics transforms applied in spectral IFS
 - \rightarrow Gaussian latitudes \Rightarrow Legendre transforms
 - $\rightarrow~$ equidistant distribution of nodes along latitudes following octahedral rule $\Rightarrow~$ Fourier transforms
- · FVM develops median-dual mesh around nodes of octahedral grid
- ⇒ finite-volume and spectral-transform IFS can operate on same quasi-uniform horizontal grid
- → Malardel et al. ECMWF Newsletter 2016, Smolarkiewicz et al. JCP 2016
- \rightarrow operational at ECMWF with HRES and ENS since March 2016
- Mesh generator and parallel data structures for FVM provided by ECMWF's Atlas framework (Deconinck et al. 2017)



Governing equations in FVM

Flux-form moist compressible Euler equations in generalised curvilinear coordinates (Smolarkiewicz, Kühnlein, Grabowski 2017; Kühnlein, Malardel, Smolarkiewicz *in prep*.):

$$\begin{split} &\frac{\partial \mathcal{G} \rho_d}{\partial t} + \nabla \cdot \left(\mathbf{v} \mathcal{G} \rho_d \right) = 0 \ , \\ &\frac{\partial \mathcal{G} \rho_d \mathbf{u}}{\partial t} + \nabla \cdot \left(\mathbf{v} \mathcal{G} \rho_d \mathbf{u} \right) = \mathcal{G} \rho_d \left[-\theta_\rho \tilde{\mathbf{G}} \nabla \varphi' + \mathbf{g} \mathcal{B} - \mathbf{f} \times \left(\mathbf{u} - \frac{\theta_\rho}{\theta_{\rho a}} \mathbf{u}_a \right) + \mathcal{M}' + \mathcal{D} + \mathcal{P}^{\mathbf{u}} \right] \ , \\ &\frac{\partial \mathcal{G} \rho_d \theta'}{\partial t} + \nabla \cdot \left(\mathbf{v} \mathcal{G} \rho_d \theta' \right) = \mathcal{G} \rho_d \left[-\tilde{\mathbf{G}}^T \mathbf{u} \cdot \nabla \theta_a + \mathcal{H} + \mathcal{P}^{\theta'} \right] \ , \\ &\frac{\partial \mathcal{G} \rho_d r_k}{\partial t} + \nabla \cdot \left(\mathbf{v} \mathcal{G} \rho_d r_k \right) = \mathcal{G} \rho_d \left[\mathcal{D}^{r_k} + \mathcal{P}^{r_k} \right] \qquad \text{where} \quad r_k = r_v, r_c, r_r, r_i, r_s \ , \\ &\varphi' = c_{pd} \left[\left(\frac{R_d}{\rho_0} \rho_d \theta \left(1 + r_v / \varepsilon \right) \right)^{R_d/c_{vd}} - \pi_a \right] \ , \end{split}$$

with:

$$\begin{split} \mathbf{v} &= \widetilde{\mathbf{G}}^T \mathbf{u} \;, \qquad \theta_\rho = \frac{\theta \left(1 + r_v / \varepsilon\right)}{\left(1 + r_t\right)} \;, \qquad \varepsilon = \frac{R_d}{R_v} \;, \qquad \theta' = \theta - \theta_{\mathfrak{s}} \;, \\ \mathcal{B} &= 1 - \frac{\theta_\rho}{\theta_{\rho\mathfrak{s}}} = 1 - \frac{\eta_{\theta\rho}}{\theta_{\rho\mathfrak{s}}} \left(\theta_{\mathfrak{s}} + \theta'\right) \;, \qquad \eta_{\theta\rho} \equiv \frac{1 + r_v / \varepsilon}{1 + r_t} \;, \qquad r_t = \sum_k r_k \;, \end{split}$$



Dry baroclinic instability (Ullrich et al. 2014) with FVM and spectral-transform IFS (ST):



 Finite-volume solutions achieve accuracy of established spectral-transform IFS for planetary-scale baroclinic instablility







Finite-volume and spectral-transform solutions in IFS

Moist baroclinic instability with FVM and spectral-transform IFS (ST) with large-scale condensation and diagnostic precipitation:





 Finite-volume solutions achieve accuracy of established spectral-transform IFS for moist flows



Finite-volume and spectral-transform solutions in IFS

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Tropical cyclone simulations with FV and ST approaches in IFS

Tropical cyclone simulations with coupling to parametrisations for large-scale condensation with diagnostic rain, surface fluxes and PBL diffusion (Reed and Jablonowski 2011) on O640/L60:





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DCMIP2016: A Review of Non-hydrostatic Dynamical Core Design and Intercomparison of Participating Models

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FVM and MPAS results for stratified flow past Schär mountain on a reduced-radius planet after 2 h (left: potential temperature perturbation θ' [K], right: vertical velocity w [m/s], lon-height section at lat=0)







Mesoscale convective storm on reduced-size planet

Supercell evolution (0.5, 1, 1.5, 2h) with FVM (left) and MPAS (right) at \approx 0.5 km grid spacing (cf. Klemp et al. 2015):



Mesoscale convective storm on reduced-size planet

Supercell for grid spacings (4, 2, 1, 0.5 km) with FVM (left) and MPAS (right) after 2 h of simulation (cf. Klemp et al. 2015):







Further reading:

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