

Behaviour of Cloud Models as seen from Asymptotic Analysis

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Juliane Rosemeier Peter Spichtinger

Institute for Atmospheric Physics (IPA)
Johannes Gutenberg University (JGU)
Mainz, Germany

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- 1 Motivation
- 2 General Cloud Model
- 3 Asymptotics
- 4 FP Analysis
- 5 Conclusion

- ▶ Clouds described by ordinary differential equations
- ▶ Derivation from first principles not possible
- ▶ Many different parameterisations
- ▶ Contain nonlinear terms
- ▶ Not clear if long term behaviour is at least similar

How can we compare the behaviour of different cloud models?

Ideas:

- ▶ Asymptotic analysis
- ▶ Theory of dynamical systems

Assumption:

- ▶ Parcel model
- ▶ Constant environmental conditions
 - ▶ temperature, pressure, supersaturation

Modeled quantities:

- ▶ Distinguish between cloud droplets and rain drops
- ▶ Cloud droplets: small, do not fall
- ▶ Rain drops: large, fall due to gravity

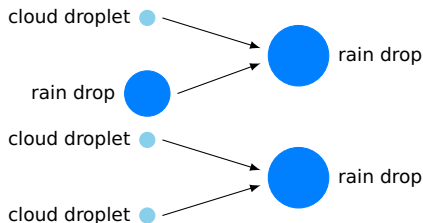
Mass concentrations:

$$q_c = \frac{\text{mass of cloud droplets}}{\text{mass of dry air}}$$

$$q_r = \frac{\text{mass of rain drops}}{\text{mass of dry air}}$$

Modeled cloud processes:

- ▶ Accretion: collisional process
- ▶ Autoconversion: collisional process
- ▶ Condensation: water vapour \leftrightarrow cloud droplets
- ▶ Evaporation: rain drops \leftrightarrow water vapour
- ▶ Sedimentation: rain drops leave air parcel due to gravity



Special treatment:

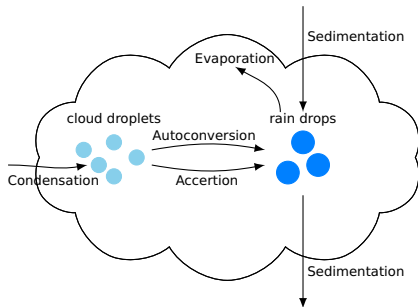
- ▶ Condensation - explicit parameterisation, no saturation adjustment
- ▶ Sedimentation



Model:

$$\frac{dq'_c}{dt'} = C - A_1 - A_2$$

$$\frac{dq'_r}{dt'} = A_1 + A_2 - E + D$$



Cloud Processes:

- ▶ Accretion: $A_1 = a'_1 q'^{\beta_c} q'^{\beta_r}$
- ▶ Autoconversion: $A_2 = a'_2 q'^{\gamma_c}$
- ▶ Condensation: $C = c' S q'_c$
- ▶ Evaporation: $E = e'_1 S q_r'^{\delta_{r1}} + e'_2 S q_r'^{\delta_{r2}}$
- ▶ Sedimentation: $D = -s' q_r'^{\zeta_r} + B'$

→ Cloud models differ by choice of parameters

Three cloud models

- ▶ Wacker: idealized model published by U. Wacker (1992)
- ▶ COSMO: warm rain scheme incorporated in COSMO model
- ▶ IFS: warm rain scheme incorporated in IFS model

Linear coefficients pressure $p = 10^5$ Pa, temperature $T = 300$ K

| | accretion | autoconversion | evaporation | | sedimentation | |
|--------|-----------|----------------------|----------------------|----------------------|----------------------|-----------|
| | a'_1 | a'_2 | e'_1 | e'_2 | s' | B' |
| Wacker | 7.5 | 10^{-4} | 0 | 0 | $3.88 \cdot 10^{-3}$ | 10^{-7} |
| COSMO | 10^{-3} | 1.96 | $3.16 \cdot 10^{-5}$ | $2.96 \cdot 10^{-4}$ | $1.29 \cdot 10^{-2}$ | 10^{-7} |
| IFS | 134 | $7.45 \cdot 10^{-2}$ | $1.79 \cdot 10^{-2}$ | $4.47 \cdot 10^{-4}$ | $4 \cdot 10^{-3}$ | 10^{-7} |

Exponents

| | accretion | | autoconversion | evaporation | | sedimentation |
|--------|-----------|---------------|----------------|---------------|-----------------|---------------|
| | β_c | β_r | γ_c | δ_{r1} | δ_{r2} | ζ_r |
| Wacker | 1 | 1 | 1 | - | - | 1 |
| COSMO | 1 | $\frac{7}{8}$ | 1 | $\frac{1}{2}$ | $\frac{11}{16}$ | $\frac{9}{8}$ |
| IFS | 1.15 | 1.15 | 2.47 | 2 | 0.635 | 1 |

Can powers and linear parameters compensate each other?

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Can powers and linear parameters compensate each other?

Reference time and reference value:

$$q_{\text{ref}} = 1.0 \cdot 10^{-4} \quad t_{\text{ref}} = 1 \text{ s}$$

Transformation:

$$q_c(t) = q_{\text{ref}} q'_c(t') \quad q_r(t) = q_{\text{ref}} q'_r(t') \quad t = t_{\text{ref}} t'$$

Compute rescaled derivatives and right hand side (RHS)

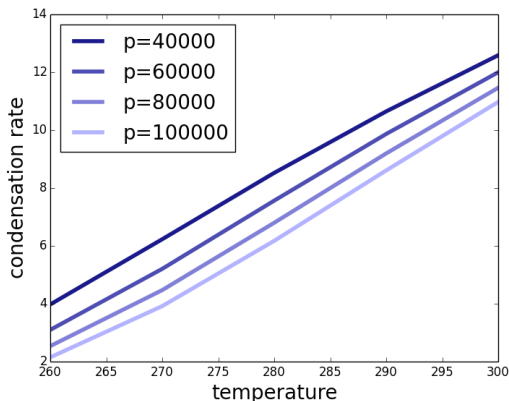
Reference system with renamed constants:

$$\begin{aligned} \frac{dq_c}{dt} &= cS q_c - a_1 q_c^{\beta_c} q_r^{\beta_r} - a_2 q_c^{\gamma_c} \\ \frac{dq_r}{dt} &= a_1 q_c^{\beta_c} q_r^{\beta_r} + a_2 q_c^{\gamma_c} - e_1 S q_r^{\delta_{r1}} - e_2 S q_r^{\delta_{r2}} - s q_r^{\zeta} + B \end{aligned}$$

Due to scaling, we can expect $q_c \approx 1, q_r \approx 1$.

1. c, a_1, a_2, e_1, e_2, B expressed in ε powers
2. Supersaturation S expressed as $S = \varepsilon^\alpha$

Caution: Some parameters depend on temperature and/or pressure!



Behaviour of models on different time scales

- ▶ Choose the time scale by another time transformation

$$\tau = \varepsilon^\omega t \quad q_c^*(\tau) = q_c(t) \quad q_r^*(\tau) = q_r(t)$$

Get a system of the form

$$\begin{aligned} \frac{dq_c^*}{d\tau} &= c^* \varepsilon^\mu q_c^* + \dots \\ \frac{dq_r^*}{d\tau} &= a_1^* \varepsilon^\nu q_c^{*\beta_c} q_r^{*\beta_r} + \dots \end{aligned}$$

with $c^*, a_1^*, \dots \in \mathcal{O}(1)$. We choose the ansatz:

$$q_c^* = q_c^{*(0)} + \varepsilon q_c^{*(1)} + \varepsilon^2 q_c^{*(2)} + \dots, \quad q_r^* = q_r^{*(0)} + \varepsilon q_r^{*(1)} + \varepsilon^2 q_r^{*(2)} + \dots$$

- ▶ Ansatz is plugged in RHS of above model
- ▶ Its derivative on the left hand side
- ▶ Order the ε powers

- ▶ Supersaturation corresponding to $\alpha = 3$ ($S = 0.1\%$)
- ▶ Time scale of $\tau = \varepsilon^3 t$ (1000 seconds)

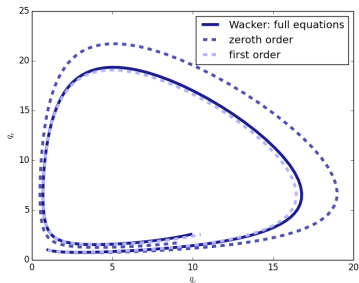
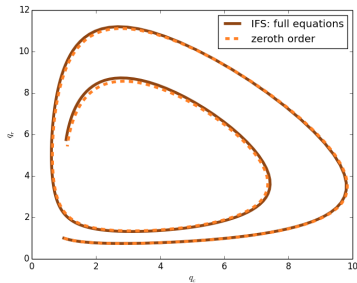
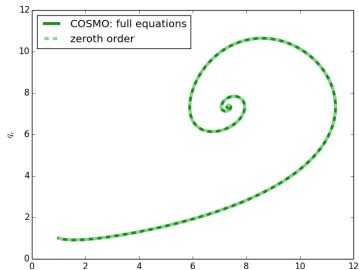
Order of processes

| | Wacker | COSMO | IFS |
|----------------|----------------------------|------------------------------|------------------------------|
| accretion | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ |
| autoconversion | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(1)$ | $\mathcal{O}(\varepsilon^3)$ |
| condensation | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ |
| evaporation | - | $\mathcal{O}(\varepsilon^3)$ | $\mathcal{O}(\varepsilon^2)$ |
| sedimentation | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ |

- ▶ Accretion, condensation, sedimentation are comparable
- ▶ Autoconversion varies with model
- ▶ Evaporation of minor importance

| | leading order equation |
|--------|--|
| Wacker | $\dot{q}_c^{(0)} = cq_c^{(0)} - a_1 q_c^{(0)} q_r^{(0)}$ $\dot{q}_r^{(0)} = a_1 q_c^{(0)} q_r^{(0)} - sq_r^{(0)} + B$ |
| COSMO | $\dot{q}_c^{(0)} = cq_c^{(0)} - a_1 q_c^{(0)} q_r^{(0)\frac{7}{8}} - a_2 q_c^{(0)}$ $\dot{q}_r^{(0)} = a_1 q_c^{(0)} q_r^{(0)\frac{7}{8}} + a_2 q_c^{(0)} - sq_r^{(0)\frac{9}{8}} + B$ |
| IFS | $\dot{q}_c^{(0)} = cq_c^{(0)} - a_1 q_c^{(0)1.15} q_r^{(0)1.15}$ $\dot{q}_r^{(0)} = a_1 q_c^{(0)1.15} q_r^{(0)1.15} - sq_r^{(0)} + B$ |

- ▶ Condensation, accretion, sedimentation on same time scale
- ▶ Discrepancy concerning time scale of autoconversion
- ▶ Predator-prey system with constant forcing



- First order approximation:

$$q_c \approx q_c^{(0)} + \varepsilon q_c^{(1)}$$

$$q_r \approx q_r^{(0)} + \varepsilon q_r^{(1)}$$

- First order approximation of Wacker model is bad for large time scales

- ▶ Keep supersaturation corresponding to $\alpha = 3$ ($S = 0.1\%$)
- ▶ Change time scale and choose $\tau = \varepsilon^2 t$ (100 seconds)

| | Wacker | COSMO | IFS |
|----------------|------------------------------|------------------------------|------------------------------|
| accretion | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ |
| autoconversion | $\mathcal{O}(\varepsilon^2)$ | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon^4)$ |
| condensation | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ |
| evaporation | - | $\mathcal{O}(\varepsilon^4)$ | $\mathcal{O}(\varepsilon^3)$ |
| sedimentation | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ |

- ▶ System scales linearly in time

| | |
|--------|--|
| | zeroth order equation |
| Wacker | $\dot{q}_c^{(0)} = 0$ $\dot{q}_r^{(0)} = 0$ |
| COSMO | $\dot{q}_c^{(0)} = 0$ $\dot{q}_r^{(0)} = 0$ |
| IFS | $\dot{q}_c^{(0)} = 0$ $\dot{q}_r^{(0)} = 0$ |

- ▶ Constant solutions in leading order
- Initial value is reproduced
- ▶ Time scale small, cloud microphysics has not yet started
- ▶ No valid long term solution

| | |
|--------|--|
| | first order equation |
| Wacker | $\dot{q}_c^{(1)} = cq_c^{(0)} - a_1q_c^{(0)}q_r^{(0)}$ $\dot{q}_r^{(1)} = a_1q_c^{(0)}q_r^{(0)} - sq_r^{(0)} + B$ |
| COSMO | $\dot{q}_c^{(1)} = cq_c^{(0)} - a_1q_c^{(0)}q_r^{(0)\frac{7}{8}} - a_2q_c^{(0)}$ $\dot{q}_r^{(1)} = a_1q_c^{(0)}q_r^{(0)\frac{7}{8}} + a_2q_c^{(0)} - sq_r^{(0)\frac{9}{8}} + B$ |
| IFS | $\dot{q}_c^{(1)} = cq_c^{(0)} - a_1q_c^{(0)1.15}q_r^{(0)1.15}$ $\dot{q}_r^{(1)} = a_1q_c^{(0)1.15}q_r^{(0)1.15} - sq_r^{(0)} + B$ |

- ▶ Changes in first order
- ▶ Structure of predator-prey system with constant forcing, **but** $q_c^{(0)} \equiv \text{const}, q_r^{(0)} \equiv \text{const}$
- ▶ Polynomial of degree 1 as first order approximation

- ▶ Increase supersaturation to $\alpha = 2$ ($S = 1\%$)
- ▶ Keep small time scale of $\tau = \varepsilon^2 t$ (100 seconds)

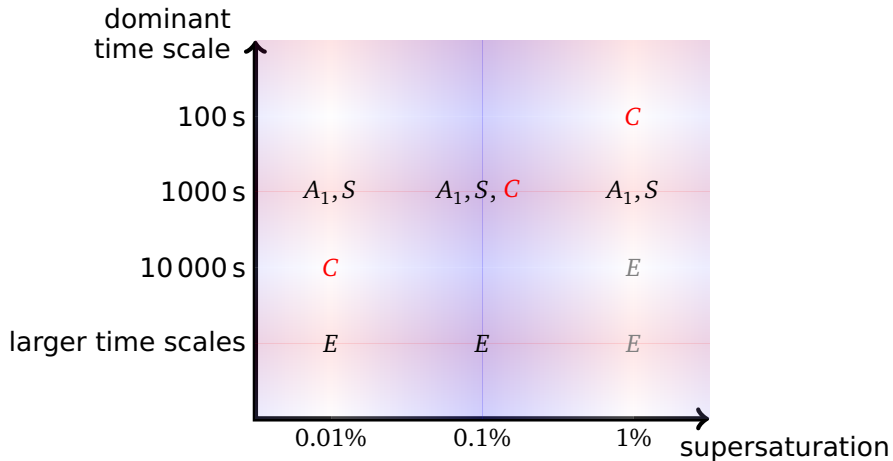
| | Wacker | COSMO | IFS |
|----------------|------------------------------|------------------------------|------------------------------|
| accretion | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ |
| autoconversion | $\mathcal{O}(\varepsilon^2)$ | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon^4)$ |
| condensation | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ | $\mathcal{O}(1)$ |
| evaporation | - | $\mathcal{O}(\varepsilon^3)$ | $\mathcal{O}(\varepsilon^2)$ |
| sedimentation | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ | $\mathcal{O}(\varepsilon)$ |

- ▶ Condensation and evaporation scale linearly in S
- ▶ Condensation dominates all processes

| | |
|--------|---|
| | leading order equation |
| Wacker | $\dot{q}_c^{(0)} = cq_c^{(0)}$ $\dot{q}_r^{(0)} = 0$ |
| COSMO | $\dot{q}_c^{(0)} = cq_c^{(0)}$ $\dot{q}_r^{(0)} = 0$ |
| IFS | $\dot{q}_c^{(0)} = cq_c^{(0)}$ $\dot{q}_r^{(0)} = 0$ |

- ▶ Condensation is the only process with an impact on the chosen time scale
 - ▶ Constant solution for q_r
 - ▶ Exponential solution for q_c
- Not realistic for large time

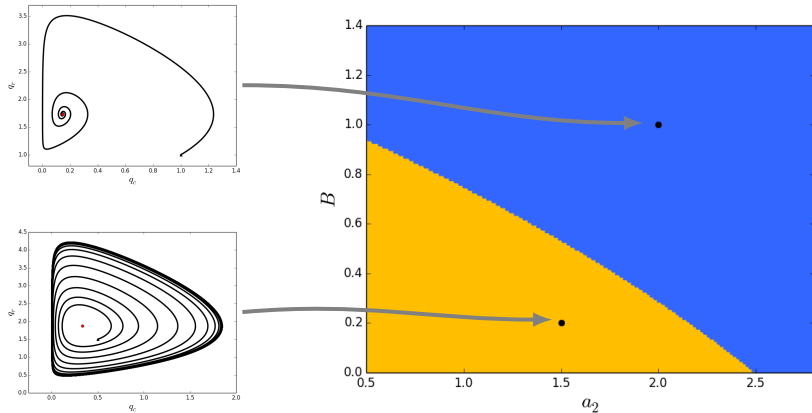
impact of A_2
depends on model



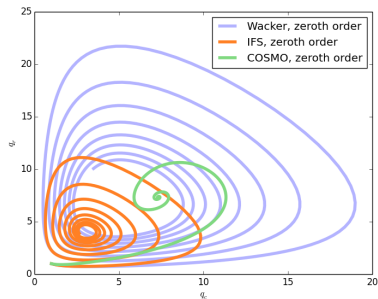
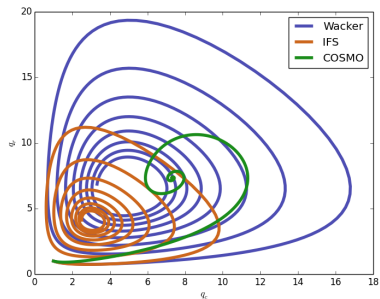
- ▶ Assumption: clouds are almost in dynamic equilibrium
- Fixed points of cloud models represent cloud
 - ▶ Usually one fixed point for $q_c = 0$
 - ▶ Wacker: $q_c = 0, q_r = 3.88$
 - ▶ Unstable if supersaturation is large enough
 - Interpretation: rain falls through, cloud develops (unstable)
 - ▶ Usually one further fixed point relevant for describing clouds
 - ▶ Wacker: $q_c = 4.87, q_r = 6.53$
 - ▶ Stability behaviour depends on parameters
 - ▶ Possibly many fixed points in general cloud model (due to modeling with fractional powers)

Stable fixed points represent the cloud as $t \rightarrow +\infty$.
Asymptotic analysis enables to look at specific time scales.

- ▶ Bifurcations can be detected in the general cloud model
 - ▶ Example below: all parameters except B and a_2 are fixed
- Autoconversion important for bifurcating behaviour



- ▶ Consider relevant fixed point $q_c \neq 0$, $S = 0.1\%$, time scale 1000 s
- ▶ For nondimensionalized Wacker, COSMO, IFS:
 - stable fixed points
 - ▶ Different time scales for approaching fixed point
- Wacker, IFS close to bifurcation
- Possible periodic feedback to coupled systems



- ▶ Models used operationally
- ▶ Both models have a fixed points of order 10^{-4}
- ▶ COSMO reaches fixed point on time scale of 1000s
- ▶ IFS reaches fixed point on larger time scale
- Many cycles around fixed point
- Possible periodic feedback to coupled system
- ▶ Zeroth order is good approximation for considered case

- ▶ Two options for adjusting the leading order system:
 - ▶ time
 - ▶ supersaturation
- ▶ All processes scale linearly in time
- ▶ Condensation, evaporation scale linearly with supersaturation
- ▶ Time scale of minutes for realistic supersaturations
- ▶ Similar models with regard to time scale of cloud processes
- ▶ However: Impact of autoconversion depends on model

- ▶ One fixed point for $q_c = 0$
 - ▶ Unstable due to supersaturation
- ▶ Usually one other relevant fixed point describing clouds
- ▶ Stable for considered cloud models
- ▶ In general stability behaviour depends on parameters
 - ▶ Bifurcations possible
 - ▶ Also determined by autoconversion
 - ▶ Periodic and almost periodic solutions possible
 - ▶ In operationally used cloud models: non periodic and almost periodic cloud models
- ▶ Can resonances occur?

Thank you for your attention!

Mitgedacht

Hier stehen die Antworten auf die Fragen.