# Entropy production due to subgrid-scale thermal fluxes with application to breaking gravity waves

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## dissipation in nature ↔ dissipation in modeling



### dissipation = temperature \* internal entropy production

resolved scales = reversible enery transformations, forth and back

unresolved scales = resolved kinetic oder internal energy are irreversibly converted into internal energy (=dissipation)





### work and heat



## **Consequences of turbulence averaging**

$$\frac{\partial}{\partial t}\,\bar{\rho}c_{\nu}\hat{T} = -\nabla\cdot(c_{\nu}\bar{\rho}\hat{\nu}\hat{T}) - \bar{p}\,\nabla\cdot\hat{\nu} - \nabla\cdot\left(\overline{c_{\nu}\rho\nu''T''}\right) - (\overline{p\nabla\cdot\nu} - \bar{p}\,\nabla\cdot\hat{\nu}) + \varepsilon_{tfr}$$



$$\frac{\partial}{\partial t}\,\bar{\rho}c_{\nu}\hat{T} = -c_{p}\overline{\Pi}\nabla\cdot\left(\bar{\rho}\widehat{\nu}\widehat{\theta}\right) - \nabla\cdot\left(\overline{c_{\nu}\rho\nu''T''}\right) + \nabla\cdot\left(\overline{c_{\nu}\rho\nu''T''}\right) - c_{p}\overline{\Pi}\nabla\cdot\left(\overline{\rho\nu''\theta''}\right) + \varepsilon_{tfr}$$

Sole approximation:  $\Pi'/\overline{\Pi} \ll 1$ This approximation is common.

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$$= \left(\frac{p}{p_0}\right)^{R/c_p} \qquad T = \Pi \theta$$

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## θ is diffused: internal entropy production positive

$$\frac{\partial}{\partial t}\,\bar{\rho}c_{\nu}\hat{T} = -c_{p}\overline{\Pi}\nabla\cdot\left(\bar{\rho}\widehat{\boldsymbol{\nu}}\widehat{\boldsymbol{\theta}}\right) - \frac{c_{p}\overline{\Pi}\nabla\cdot\left(\bar{\rho}\boldsymbol{\nu}^{"}\boldsymbol{\theta}^{"}\right) + \varepsilon_{tfr}}{\sqrt{\rho}\hat{T}\frac{\hat{d}}{dt}\hat{s}}$$

Second law of thermodynamics



gradient approach:

$$\overline{\rho \boldsymbol{v}^{"} \boldsymbol{\theta}^{"}} = -\rho \underline{\boldsymbol{K}^{\theta}} \cdot \boldsymbol{\nabla} \widehat{\boldsymbol{\theta}}$$

$$\sigma_{\theta} = \frac{c_p \bar{\rho}}{\hat{\theta}^2} K_{ii}^{\theta} (\partial_i \hat{\theta})^2 \ge 0$$

always positive, regardless of stratification

dissipation by  $\theta$ -diffusion:  $\varepsilon_{\theta} = \hat{T}\sigma_{\theta}$ 



## **BUT: Inspect energy conversions!**

$$\frac{\partial}{\partial t} \bar{\rho} c_{v} \hat{T} = -c_{p} \overline{\Pi} \nabla \cdot \left( \bar{\rho} \widehat{v} \widehat{\theta} \right) - c_{p} \overline{\Pi} \nabla \cdot \left( \bar{\rho} v^{"} \overline{\theta}^{"} \right) + \varepsilon_{tfr} \qquad \psi \nabla \cdot f + f \cdot \nabla \psi = \nabla \cdot \left( \psi f \right)$$

$$\frac{\partial}{\partial t} \bar{\rho} c_{v} \hat{T} = -c_{p} \overline{\Pi} \nabla \cdot \left( \bar{\rho} \widehat{v} \widehat{\theta} \right) - \nabla \cdot \left( c_{p} \overline{\Pi} \overline{\rho} v^{"} \overline{\theta}^{"} \right) + c_{p} \overline{\rho} \overline{v}^{"} \overline{\theta}^{"} \cdot \nabla \overline{\Pi} + \varepsilon_{tfr}$$

Energy exchange with kinetic energy has not been inspected thorougly enough!

Consider only vertical fluxes

$$c_{p}\overline{\rho w''\theta''}\partial_{z}\overline{\Pi} = -c_{p}\overline{\rho}K^{\theta}\partial_{z}\theta\left(-\frac{g}{c_{p}\widehat{\theta}}\right) = \overline{\rho}K^{\theta}N^{2}$$
$$\overline{\rho}K^{\theta}N^{2} < 0$$

### $\bar{\rho}K^{\theta}N^2 > 0$

- gain of internal energy
- entropy production meaningful
- loss of resolved kinetic energy
- a force must represent this kinetic energy loss in the momentum equation

- loss of internal energy
- entropy production meaningless!
- making it meaningful must prevent the gain of resolved kinetic energy
- instead, TKE is generated, but TKE is indistinguishable from internal energy
- the traditional approach (grey box) is safe

### **Case distinction necessary!**



## Case distinction: N<sup>2</sup><0

omit resolved energy conversion, only applicable to N<sup>2</sup><0</li>

$$\begin{aligned} \frac{\partial}{\partial t} \ \bar{\rho} c_{v} \hat{T} &= -c_{p} \overline{\Pi} \nabla \cdot \left( \bar{\rho} \widehat{v} \widehat{\theta} \right) - \nabla \cdot \left( c_{p} \overline{\Pi} \rho v^{"} \theta^{"} \right) + \varepsilon_{tfr}}{\rho \overline{T}} \int_{d} \hat{d}_{t} \hat{s} \\ \bar{\rho} \frac{\partial}{\partial t} \hat{s} &= -\nabla \cdot \left( \frac{c_{p} \overline{\rho} v^{"} \theta^{"}}{\widehat{\theta}} \right) - \underbrace{\frac{c_{p} \overline{\Pi} \overline{\rho} v^{"} \theta^{"}}{\widehat{T}^{2}} \cdot \nabla \hat{T} + \varepsilon_{tfr} / \hat{T}}_{\text{export / import}} & \text{internal entropy production} \end{aligned}$$

$$gradient apporoach: c_{p} \overline{\Pi} \overline{\rho} v^{"} \theta^{"} &= -c_{p} \overline{\rho} \underline{K}^{T} \cdot \nabla \hat{T} & c_{p} \overline{\Pi} \overline{\rho} w^{"} \theta^{"} &= -c_{p} \overline{\rho} \overline{\Pi} K^{\theta} (\partial_{z} \hat{\theta} - \gamma) \int_{d} \hat{T} \geq 0 & \sigma_{T} &= \frac{c_{p} \overline{\rho}}{\widehat{T}^{2}} \overline{\Pi} K^{\theta} (\partial_{z} \hat{\theta} - \gamma) \partial_{z} \widehat{T} \geq 0 \end{aligned}$$

Formally, this is a temperature diffusion = subscale heat flux.

For unstable stratification  $\partial_z \hat{\theta}$  and  $\partial_z \hat{T}$  are parallel.

Dissipation by T-diffusion:  $\varepsilon_T = \hat{T}\sigma_T$ 

Contradiction to 2nd law, if applied in case of N<sup>2</sup>>0



### Case distinction: N<sup>2</sup>>0

$$\frac{\partial}{\partial t} \bar{\rho} c_{v} \hat{T} = -c_{p} \overline{\Pi} \nabla \cdot (\bar{\rho} \widehat{v} \widehat{\theta}) - c_{p} \overline{\Pi} \nabla \cdot (\bar{\rho} v^{"} \theta^{"}) + \varepsilon_{tfr} \qquad \varepsilon_{tfr} = -\bar{\rho} v^{"} v^{"} \cdot \nabla \widehat{v} \ge 0$$

$$\psi \nabla \cdot f + f \cdot \nabla \psi = \nabla \cdot (\psi f)$$

$$\frac{\partial}{\partial t} \bar{\rho} \left( \frac{\widehat{v}^{2}}{2} + \Phi \right) = -c_{p} \bar{\rho} \widehat{v} \widehat{\theta} \cdot \nabla \overline{\Pi} - c_{p} \overline{\rho} v^{"} \theta^{"} \cdot \nabla \overline{\Pi} - \widehat{v} \cdot \nabla \cdot \overline{\rho} v^{"} v^{"} - \nabla \cdot \left( \bar{\rho} \widehat{v} \left( \frac{\widehat{v}^{2}}{2} + \Phi \right) \right)$$

#### Which momentum equation belongs to kinetic energy equation?

Consider only vertical direction.

$$\begin{split} &\frac{\partial}{\partial t}\,\widehat{w} = -g - c_p \widehat{\theta} \partial_z \overline{\Pi} - \frac{c_p \frac{\overline{\rho w'' \theta''}}{\overline{\rho \hat{w}}} \partial_z \overline{\Pi}}{\frac{\partial}{\partial t}\,\widehat{w} = -g - c_p \widehat{\theta} \partial_z \overline{\Pi} - \frac{K^{\theta} N^2}{\widehat{w}}}. \end{split}$$

$$\frac{\partial}{\partial t}\,\widehat{w} = -g - c_p \widehat{\theta} \partial_z \overline{\Pi} - R_w \widehat{w}$$

 $-c_p\hat{\theta}\partial_z\overline{\Pi} = -\frac{1}{\bar{o}}\partial_z\bar{p}$ 

- new term
- turbulent pressure gradient term
- similarity to Rayleigh damping

$$R_w = N^2 K^\theta / \widehat{w}^2$$

- diffusion coefficient must prevent singularity
- new term leads to downward turbulent θ-flux.



### Case distinction: N<sup>2</sup>>0

$$\frac{\partial}{\partial t}\,\widehat{w} = -g - c_p \hat{\theta} \partial_z \overline{\Pi} - R_w \widehat{w}$$

$$R_w = N^2 K^\theta / \widehat{w}^2$$

**Hypothesis**: For shortest resolvable scales, the horizonal wind is damped by vertical diffusion as fast as the vertical wind is damped by Rayleigh damping.

$$R_w = K^m \frac{\pi^2}{(\Delta z)^2} \qquad \qquad K^\theta = K^m \frac{\pi^{(w)}}{(\Delta z)^2 N^2}$$

There is no diffusion for  $\widehat{w}^2 = 0$ .

### Consider isentropes of a breaking gravity wave



## Examplary 2-d modeling with ICON-IAP

ICON-IAP model with hexagonal mesh (QJRMS,2013)



 $\Delta z = 250 \text{ m}, \Delta x = 2 \text{ km}, \Delta t = 3 \text{ s},$ H=120 km, L=1200 km, T=32 h

- $K^m$  as in Holtslag und Boville (1993)
- initial profile as in Chun and Kim (2008)

(a) (b) 105 90. 90. 75-75z (km) 60-60-45-45. 30. 30 15-15. -75 -50 -25 0 25 50 75 250 150 200 300 U (m s<sup>-1</sup>) T (K)

**Figure 1.** The basic-state (a) zonal wind and (b) temperature used for the numerical simulations. These are the July mean values at  $35^{\circ}$ N from the CIRA climate data.

• gravity wave generator as in Durran (1999), ceases after 16 hours



## w, $\theta$ and u in breaking region



- isentropes have local minimum at w = 0
- gravity wave breaks near the critical level:  $m^2 \rightarrow \infty$
- isentropes overturn

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vertical wind shear is large



### **Downward directed θ fluxes**

 $E_{new} = \rho K^m w^2 \pi^2 / (\Delta z)^2$ 

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 $E_{old} = \rho K^m N^2$ 





### 3 setups

#### EXP 1: state of the art, inconsistent for N<sup>2</sup>>0



#### EXP 4: nothing for N<sup>2</sup>>0

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#### After a long time....

#### EXP 2: entropically consistent for N<sup>2</sup>>0



Further experiments (not shown):

- forcing in w-eq. is omitted, but θ-flux is retained in θ-eq,
- typical numerical off-centering in the implicit solver for w is used, results are very similar to exp 2.



A. Gassmann: Entropy production by subgrid  $\theta$ -fluxes and gravity wave breaking

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## Relative static stability N<sup>2</sup>/N<sup>2</sup><sub>iso</sub>



EXP 1: state of the are, inconsistent for N<sup>2</sup>>0

### EXP 2: entropically consistent for N<sup>2</sup>>0

### EXP 4: nothing for N<sup>2</sup>>0

 $N_{iso}^2 = g^2/(cpT)$ N<sup>2</sup> for isothermal stratification



## **Temperature profiles**



Lidar-measurements (Liu and Meriwether, 2004)

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Experiments 1,2,4

- sharp maxima, peaks
- sometimes overadiabatic stratification
- no parabolic shape



### Horizontal mean und variability of N<sup>2</sup>/N<sup>2</sup><sub>iso</sub>



## Dissipations rates $\varepsilon_{\theta}, \varepsilon_T, \varepsilon_{tfr}$



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 $\uparrow$  frictional dissipation is 10 times larger than thermal dissipation

← thermal dissipation

#### upper picture:

- $\partial_z \theta \partial_z T$  changes sign at isothermal stratification
- inversion layer (  $\partial_z T > 0$  ) has positive dissipation, but the physical process is wrong
- there should not be any qualitative difference between less and more stable stratification than isothermal stratification, if the stratification is stable

## Conclusion

**Case distinction** for subscale parameterization:

stable N2>0 and unstable N2<0

**Energy conversion** between resolved kinetic and internal energy is necessary for N<sup>2</sup>>0.

This requires a new term in the vertical momentum equation.

The friction term converts likewise kinetic energy into internal energy.

But this process is described by tensor fluxes and not by vector fluxes.

